

STARFlow-V: End-to-End Video Generative Modeling with Normalizing Flows

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Normalizing flows (NFs) are end-to-end likelihood-based generative models for continuous data, and have recently regained attention with encouraging progress on image generation. Yet in the video generation domain, where spatiotemporal complexity and computational cost are substantially higher, state-of-the-art systems almost exclusively rely on diffusion-based models. In this work, we revisit this design space by presenting STARFlow-V, a normalizing flow-based video generator with substantial benefits such as end-to-end learning, robust causal prediction, and native likelihood estimation. Building upon the recently proposed STARFlow, STARFlow-V operates in the spatiotemporal latent space with a global-local architecture which restricts causal dependencies to a global latent space while preserving rich local withinframe interactions. This eases error accumulation over time, a common pitfall of standard autoregressive diffusion model generation. Additionally, we propose flow-score matching, which equips the model with a light-weight causal denoiser to improve the video generation consistency in an autoregressive fashion. To improve the sampling efficiency, STARFlow-V employs a video-aware Jacobi iteration scheme that recasts inner updates as parallelizable iterations without breaking causality. Thanks to the invertible structure, the same model can natively support text-to-video, image-to-video as well as video-to-video generation tasks. Empirically, STARFlow-V achieves strong visual fidelity and temporal consistency with practical sampling throughput relative to diffusion-based baselines. These results present the first evidence, to our knowledge, that NFs are capable of high-quality autoregressive video generation, establishing them as a promising research direction for building world models.

Code: https://github.com/apple/ml-starflow
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1 Introduction

Deep generative modeling has advanced rapidly with breakthroughs across language (Achiam et al., 2023; OpenAI, 2024a), images (Podell et al., 2023; Batifol et al., 2025; Wu et al., 2025), and videos (OpenAI, 2024b; Wan et al., 2025; DeepMind, 2025) domains. Among these modalities, video generation is uniquely demanding: beyond high perceptual quality, models must capture rich spatiotemporal structure, remain robust over long horizons, and often operate under causal constraints for streaming and interactive use. Such capabilities are central not only to creative media (Ye et al., 2025; Yuan et al., 2025), but also to emerging world models for gaming, simulation and embodied AI (Ha and Schmidhuber, 2018; Yang et al., 2023; Hu et al., 2023; Google DeepMind, 2024; Hafner et al., 2025).

Recent scaling of data, model capacity, and compute has pushed video generation to new levels of fidelity (Yang et al., 2025; Kong et al., 2024; Kondratyuk et al., 2024; Yu et al., 2024; Wan et al., 2025; Seawead et al., 2025; Gao et al., 2025). In this space, diffusion-based approaches (Ho et al., 2020; Rombach et al., 2022; Peebles and Xie, 2023; Lipman et al., 2023; Esser et al., 2024) have emerged as the dominant

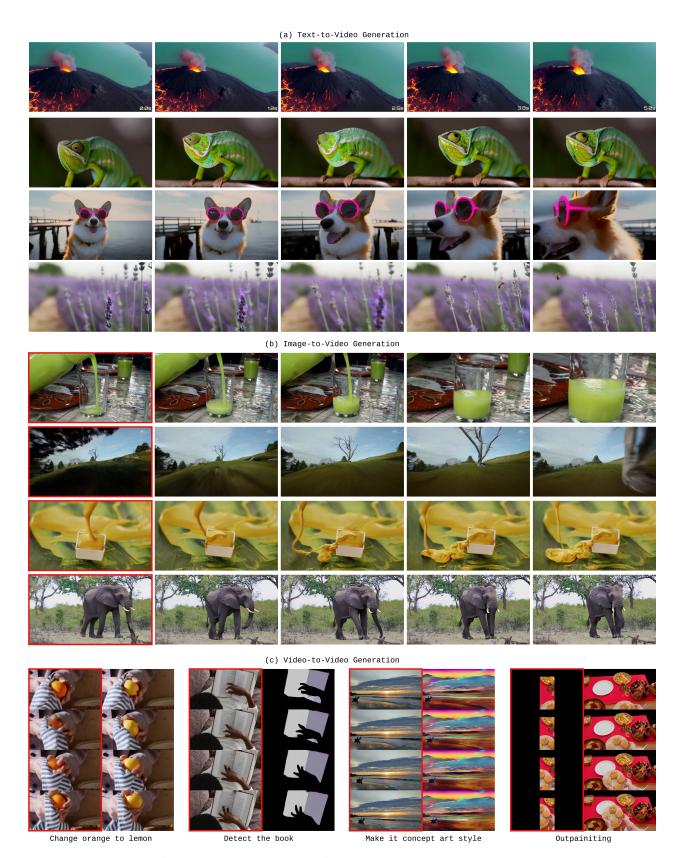


Figure 1 Samples from STARFlow-V across three tasks. All videos are 480 p at 16 fps. Red boxes mark the conditioning inputs. The same autoregressive architecture is used for all tasks with no task-specific modifications. **Please find more generated videos and comparisons in the released code https://github.com/apple/ml-starflow.**

backbone for text- and image-conditioned video synthesis, thanks to their strong empirical performance and flexible conditioning mechanisms. Standard diffusion models are trained by corrupting frames with noise drawn from a schedule and learning a denoiser that inverts this process one step at a time, which leads to an iterative sampling procedure at inference. For offline generation this formulation works well, but the parallel denoising of multiple frames is inherently non-causal: future frames can influence earlier ones, making it less natural to apply in streaming or interactive settings that require strictly causal rollouts. Causally conditioned and sequential diffusion variants (Chen et al., 2024a; Huang et al., 2025) mitigate some of these issues, but still inherit the need to simulate noise at different timesteps and frames during training and can exhibit train—test mismatch during long-horizon autoregressive generation.

In parallel, normalizing flows (NFs) (Rezende and Mohamed, 2015; Dinh et al., 2014, 2016) offer a distinct, likelihood-based alternative. NFs are continuous end-to-end generative models that provide exact log-likelihood evaluation, non-iterative sampling, and native support for invertible feature mappings. After an initial wave of work (Dinh et al., 2016; Kingma and Dhariwal, 2018), they received relatively less attention compared to diffusion models, but have recently regained interest with encouraging progress on image generation (Zhai et al.; Gu et al., 2025; Zheng et al., 2025). In particular, STARFlow (Gu et al., 2025) shows that parameterizing an "autoregressive normalizing flow" with a Transformer and operating in a latent space allows flows to scale competitively in the high-resolution image domain. Yet, in the video domain—where complexity and computational cost are substantially higher—state-of-the-art systems almost exclusively rely on diffusion, and it remains unclear whether NFs can be practical for video.

In this work, we revisit this design space and introduce STARFlow-V, a normalizing-flow-based video generator that combines end-to-end training with causal, likelihood-based modeling. Building on STARFlow (Gu et al., 2025), STARFlow-V operates in a spatiotemporal latent space with a global-local architecture: a compact global latent sequence carries long-range temporal context, while local latent blocks preserve fine-grained within-frame structure. By delegating temporal reasoning to this high-level space, the model mitigates the accumulation of autoregressive errors that commonly plagues diffusion-based video generators. As observed in TARFlow (Zhai et al., 2024), training flows on slightly perturbed data with a subsequent denoising step can significantly improve robustness. Unlike existing methods (Zhai et al., 2024; Gu et al., 2025), we propose flow-score matching, which learns a lightweight causal denoiser to enhance temporal consistency in video scenarios. To further improve efficiency, STARFlow-V employs a video-aware Jacobi-style update scheme that recasts inner refinement steps as parallelizable iterations. Finally, owing to its invertible nature, the same backbone naturally supports text-to-video (T2V), image-to-video (I2V), and video-to-video (V2V) generation by simply changing the form of the conditioning signal.

Across all benchmarks, STARFlow-V attains visually coherent and temporally stable generations while maintaining practical sampling speed relative to diffusion-based models. We believe this provides **initial** evidence that NFs are capable of high-quality autoregressive video generation and potentially world models.

2 Background

2.1 Video Generative Models

Given N frames $\mathbf{x}_{1:N} = (\mathbf{x}_1, \dots, \mathbf{x}_N)$ and optional conditioning C (e.g., text, image, audio, layout, camera), video generative models seek to model the joint distribution of all frames $p(\mathbf{x}_{1:N} \mid C)$ and sample novel videos from the learned model. While earlier work explored GANs (Vondrick et al., 2016; Tulyakov et al., 2018; Skorokhodov et al., 2022), VAEs (Babaeizadeh et al., 2018; Castrejon et al., 2019; Wu et al., 2021), and discrete autoregressive models (Yan et al., 2021; Yu et al., 2024; Kondratyuk et al., 2024), the field has largely converged on diffusion-based methods Ho et al. (2022c,a). Spurred by the release of Sora (Brooks et al., 2024), DiT-style approaches (Peebles and Xie, 2023) have shown strong generalization at scale (Gao et al., 2025; Wan et al., 2025; DeepMind, 2025). A key distinction from prior paradigms is that training of diffusion-based models is Not End-to-End: diffusion-based models corrupt frames with noise at randomly sampled levels and train a denoiser to invert this process, optimizing an objective closely related to the lower bound of $\log p(\mathbf{x}_{1:N} \mid C)$. This setup incurs high cost—especially for video—as each update supervises only a single noise level. At inference time, one sample is generated by iteratively denoising from Gaussian noise.

Diffusion-based video generation is typically non-causal: all frames are corrupted with noise and denoised in parallel (Ho et al., 2022c). Yet many real-world applications demand causal, often interactive synthesis (e.g., online streaming, video games, robotics), where frames must be produced sequentially. Autoregressive (AR) diffusion models (Chen et al., 2024a; Song et al., 2025; Yin et al., 2025)—a line of work that combines chain-rule factorization with diffusion—aim to alleviate prior limitations by introducing asynchronous, frame-wise noise schedules during training, modeling each conditional $p(x_n \mid x_{< n})$ as a diffusion process. Despite their strengths, AR generation typically suffers from exposure bias: during training, models condition on ground-truth contexts, whereas at inference they must rely on their own (imperfect) predictions. This train—test mismatch compounds over time, degrading long-horizon video quality. The non-end-to-end nature of diffusion training further exacerbates this gap, though recent efforts such as Self-Forcing (Huang et al., 2025) seek to mitigate it via sequential post-training with distillation objectives. However, they are not readily applicable in the pre-training stage on raw video data.

2.2 Autoregressive Normalizing Flows

Normalizing flows (NFs; Rezende and Mohamed, 2015; Dinh et al., 2014, 2016; Kingma and Dhariwal, 2018; Ho et al., 2019) are likelihood-based generative models built from invertible transformations. Given a continuous input $\boldsymbol{x} \sim p_{\text{data}}, \ \boldsymbol{x} \in \mathbb{R}^D$, an NF learns a bijection $f_{\theta} : \mathbb{R}^D \to \mathbb{R}^D$ that maps data \boldsymbol{x} to latents $\boldsymbol{z} = f_{\theta}(\boldsymbol{x})$. Unlike diffusion models, NFs are trained *end-to-end* via a tractable maximum-likelihood objective derived from the change-of-variables formula:

$$\mathcal{L}_{NF}(\theta) = \mathbb{E}_{\boldsymbol{x}} \left[\log p_0(f_{\theta}(\boldsymbol{x})) + \log |\det(J_{f_{\theta}}(\boldsymbol{x}))| \right], \tag{2.1}$$

where the first term encourages mapping data to high-density regions of a simple prior p_0 (e.g., standard Gaussian), and the Jacobian term J_f accounts for the local volume change induced by f_θ , preventing collapse. Once trained, sampling is immediate via inversion: draw $z \sim p_0(z)$ and set $x = f_\theta^{-1}(z)$. Historically, however, NFs have been viewed as less competitive than diffusion models due to architectural rigidity and training instability (Dinh et al., 2016).

Recently, TARFlow (Zhai et al.) and its scalable extension, STARFlow (Gu et al., 2025), have revisited normalizing flows as next-generation backbones for generative modeling. Both methods instantiate autoregressive flows (AFs) (Kingma et al., 2016; Papamakarios et al., 2017)—NFs whose invertible transformations are parameterized autoregressively—and use causal Transformer blocks, in the style of LLMs, as their primary building units. Formally, STARFlow (Gu et al., 2025) stacks T autoregressive flow blocks with alternating directions, where each block applies an affine transform whose parameters are predicted by a causal Transformer under a (self-exclusive) causal mask m:

$$z = \left[x - \mu_{\theta}(x \odot m) \right] / \sigma_{\theta}(x \odot m), \ \sigma_{\theta}(\cdot) > 0, \tag{2.2}$$

where x, z are the input and output of each block, \odot denotes the Hadamard product. As shown in STARFlow (Gu et al., 2025), $T \ge 3$ blocks suffice for universal density modeling where masks alternate between left-to-right (\rightarrow) and right-to-left (\leftarrow) to capture bidirectional dependencies.

Despite STARFlow demonstrating competitive quality with state-of-the-art diffusion (Podell et al., 2023; Esser et al., 2024) on large-scale text-to-image tasks, evidence for normalizing flows in video generation remains sparse. To our best knowledge, the only prior NF-based video model is VideoFlow (Kumar et al., 2019), which builds on Glow (Kingma and Dhariwal, 2018) and is constrained by limited capacity, low resolution, and domain-specific settings. Compared to images, video generation is substantially more challenging for NFs due to higher spatiotemporal dimensionality. Nevertheless, we argue that normalizing flows—exemplified by STARFlow—are a natural fit for video modeling, especially in autoregressive settings.

3 STARFlow-V

We propose STARFlow-V, a novel paradigm for video generation based on normalizing flows. While inspired by STARFlow (Gu et al., 2025), STARFlow-V is not a direct port to the video domain; it introduces several architectural redesigns and algorithmic innovations tailored to spatiotemporal data. In what follows, we present the architecture and its autoregressive formulation (Section 3.1), the training procedure (Section 3.2), the inference pipeline (Section 3.3), and applications enabled by our model (Section 3.4).

3.1 Proposed Model

For a video $\boldsymbol{x} \in \mathbb{R}^{N \times H \times W \times D}$, each frame \boldsymbol{x}_n is flattened to $\mathbb{R}^{HW \times D}$, $\boldsymbol{x}_n = (\boldsymbol{x}_{n,1}, \dots, \boldsymbol{x}_{n,HW})$, and all frames are concatenated into a sequence of NHW tokens. We operate in a compressed latent space using a pretrained 3D causal VAE (Wan et al., 2025). STARFlow-V models the joint distribution $p_{\theta}(\boldsymbol{x})$ via an invertible mapping f_{θ} implemented as autoregressive flows (Equation (2.2)). Following Gu et al. (2025), we use a deep-shallow decomposition $f_{\theta} = f_D \circ f_S$, where a small stack of shallow flow blocks with alternating (left-to-right / right-to-left) masks maps \boldsymbol{x} to intermediate latents $\boldsymbol{u} = f_S(\boldsymbol{x})$, and a deep causal-Transformer flow f_D then maps \boldsymbol{u} to the prior, producing $\boldsymbol{z} = f_D(\boldsymbol{u})$. By the change-of-variables formula,

$$p_{\theta}(\boldsymbol{x}) = p_{0}(\boldsymbol{z}) \left| \det J_{f_{D}}(\boldsymbol{u}) \right| \left| \det J_{f_{S}}(\boldsymbol{x}) \right|, \tag{3.1}$$

where p_0 is a simple prior (e.g., standard Gaussian). Most capacity is allocated to the deep block f_D for semantic modeling, while the shallow stack f_S handles local reshaping. For videos, we can simply treat all frames as one long token sequence: f_D follows a left-to-right causal order over the video (causal across frames, raster order within each frame), and f_S retains the alternating masks defined above. Because f_S propagates information from future frames to past ones, this naïve design yields a *non-causal* video generator, motivating the global–local restructuring described next.

Global-Local Architecture Observing that f_D is inherently autoregressive and that f_S mainly provides local refinements, we adapt the design into a global-local structure: f_S is restricted to operate within each frame, while only f_D propagates global video context in a causal manner. More specifically, Equation (3.1) can be re-expressed as an autoregressive factorization over frames x_n :

$$p_{\theta}(\boldsymbol{x}) = \prod_{n=1}^{N} p_{\theta}(\boldsymbol{x}_n \mid \boldsymbol{x}_{< n}) = \prod_{n=1}^{N} p_D(\boldsymbol{u}_n \mid \boldsymbol{u}_{< n}) |\det J_{f_S}(\boldsymbol{x}_n)|,$$
(3.2)

where $u_n = f_S(x_n)$ denotes the local latents for frame x_n . Here, the deep block is itself an autoregressive flow, capturing both intra-frame raster ordering and inter-frame causal dependencies.

Formulating STARFlow-V in a *global-local* manner (Equation (3.2)) yields several benefits:

- (a) Universality. Equation (3.2) preserves the universal approximation guarantee of STARFlow (Gu et al., 2025): the local stack f_S still realizes per-pixel infinite Gaussian mixtures via alternating causal masks, so expressivity is not curtailed by restricting f_S to within-frame contexts.
- (b) Robustness. Intuitively, Equation (3.2) can be viewed as a continuous language model for videos: the deep-flow term $p_D(\boldsymbol{u}_n \mid \boldsymbol{u}_{\leq n})$ acts as Gaussian Next-Token Prediction (cf. the affine form in Equation (2.2)) in latent space, while the shallow flow supplies the Jacobian factor $|\det J_{f_S}(\boldsymbol{x}_n)|$, yielding a flexible density over \boldsymbol{x} . Compared to modeling \boldsymbol{x} directly (arbitrarily multimodal), the latent \boldsymbol{u} is unimodal at each step, easier to regress, and more tolerant to small prediction errors. Crucially, the sampling phase via f_D^{-1} conditions on previously generated latents rather than pixels, so data-space errors do not propagate forward, mitigating the compounding error typical of autoregressive diffusion. Unlike diffusion-style noise conditioning (Ho et al., 2022b; Chen et al., 2024a), which compromises information to gain robustness and introduces extra parameters, our mappings $\boldsymbol{u} \leftrightarrow \boldsymbol{x}$ are invertible, avoiding information loss by construction.
- (c) **End-to-End Training.** The whole model is still NF. Consequently, all parameters are trained by exact MLE via the change-of-variables objective—no per-step denoising schedule or surrogate loss—simplifying optimization and reducing train—test mismatch.
- (d) **Streamable Generation.** At inference time, f_D^{-1} samples u_n causally (token-by-token, frame-by-frame), and f_S^{-1} decodes each frame independently given u_n . This process enables causal video synthesis since later frames cannot influence earlier ones.

3.2 Revisiting Noise-Augmented Training

As observed by Zhai et al. (2024), injecting *small* noise into the data is crucial for stabilizing NF training. Concretely, we learn STARFlow-V on a σ -smoothed density $q_{\sigma}(\tilde{x}) = (p * \mathcal{N}(0, \sigma^2 I))(\tilde{x})$. A side effect is that

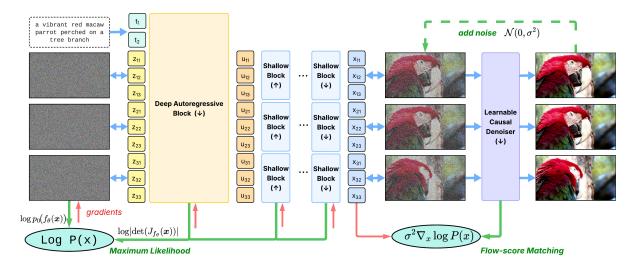


Figure 2 An illustrated pipeline of STARFlow-V which shows (1) the proposed global-local architecture; (2) joint training with the learnable denoiser with the proposed Flow-score Matching. During sampling, STARFlow-V takes the encoded text condition \boldsymbol{t} and transforms the noise \boldsymbol{z} through deep global block to intermediate features \boldsymbol{u} , followed by several local shallow blocks to produce a slightly noised video. Finally, a learnable causal denoiser refines this output into the final clean video \boldsymbol{x} .

the model naturally generates slightly noisy samples, necessitating a post-processing step to recover the clean ones. We first examined the existing options for this purpose:

- (a) **Decoder Fine-tuning** We followed STARFlow (Gu et al., 2025), adopting their strategy of fine-tuning the VAE decoder to denoise noisy latents using a GAN objective (Rombach et al., 2022). However, our preliminary experiments suggest that this approach is not readily applicable to 3D causal VAEs: under Gaussian-noised latent inputs, the decoder fails to maintain temporal consistency in the generated videos due to limited receptive fields.
- (b) **Score-based Denoising** Instead of decoder fine-tuning, TARFlow (Zhai et al., 2024) proposes to denoise using the *learned flow* itself via score-based updates. For a noisy sample $\tilde{x} \sim q_{\sigma}$, the continuity equation gives $\partial_{\sigma}\tilde{x} = -\sigma \nabla_{\tilde{x}} \log q_{\sigma}(\tilde{x})$. So for sufficiently small σ , a single Euler step yields the Tweedie estimator:

$$x \approx \tilde{x} - \sigma \partial_{\sigma} \tilde{x} = \tilde{x} + \sigma^2 \nabla_{\tilde{x}} \log q_{\sigma}(\tilde{x}).$$
 (3.3)

With normalizing flows, we replace q_{σ} by the learned density p_{θ} , and compute $\nabla_{\tilde{\boldsymbol{x}}} \log p_{\theta}(\tilde{\boldsymbol{x}})$ via automatic differentiation through the flow, which amounts to an additional forward–backward pass. However, this score-based denoising presents two issues: (1) Noisy gradients. The learned density p_{θ} is imperfect; its score $\nabla_{\tilde{\boldsymbol{x}}} \log p_{\theta}(\tilde{\boldsymbol{x}})$ often contains high-frequency noise, which manifests as bright speckle-like artifacts—especially in regions with large motion; (2) Non-causality of the score. Even if p_{θ} is modeled causally, the score $\nabla_{\tilde{\boldsymbol{x}}} \log p_{\theta}(\tilde{\boldsymbol{x}})$ is, by definition, global: the gradient at time n depends on likelihood terms involving future frames m > n. This breaks causality, undermining the promised streamable generation.

Proposed Approach: Flow-Score Matching To address these issues, we introduce a lightweight neural denoiser s_{ϕ} trained alongside the flow f_{θ} to regress the model's score:

$$\mathcal{L}_{\text{denoise}}(\phi) = \mathbb{E}_{\boldsymbol{x}, \epsilon} \| s_{\phi}(\tilde{\boldsymbol{x}}) - \sigma \nabla_{\tilde{\boldsymbol{x}}} \log p_{\theta}(\tilde{\boldsymbol{x}}) \|_{2}^{2}, \qquad \tilde{\boldsymbol{x}} = \boldsymbol{x} + \boldsymbol{\epsilon}, \ \boldsymbol{\epsilon} \sim \mathcal{N}(0, \sigma^{2} I).$$
 (3.4)

At inference, we replace the raw score in the update (cf. Equation (3.3)) with the learned denoiser s_{ϕ} . This flow-score matching (FSM) is simple yet effective. First, the smooth inductive bias of neural networks suppresses stochastic high-frequency artifacts in $\nabla_{\tilde{x}} \log p_{\theta}$. Second, we can encode causality directly in s_{ϕ} , reensuring streamable behavior. Concretely, we parameterize s_{ϕ} with a one-frame look-ahead while remaining globally causal (one-step latency)¹. We approximate the score at step p_{ϕ} by $s_{\phi}(\tilde{x}_{\leq n+1}) \approx (\sigma \nabla_{\tilde{x}} \log p_{\theta}(\tilde{x}))_{\pi}$.

¹Strictly causal ($\leq n$) fails as temporal differences are pivotal to determining the denoising direction.

Model	Total	Quality	Semantic	Aesthetic	Object	Multi Obj.	Human	Spatial	Scene
Closed-source models									
Gen-3 (Germanidis, 2024)	82.32	84.11	75.17	63.34	87.81	53.64	96.40	65.09	54.57
Veo3 (Google DeepMind, 2025)	85.06	85.70	82.49	63.81	93.89	82.20	99.40	84.26	57.43
Diffusion models									
OpenSora-v1.1 (Zheng et al., 2024)	75.66	77.74	67.36	50.12	86.76	40.97	84.20	52.47	38.63
CogVideoX (Yang et al., 2024)	80.91	82.18	75.83	60.82	83.37	62.63	98.00	69.90	51.14
HunyuanVideo (Kong et al., 2024)	83.24	85.09	75.82	60.36	86.10	68.55	94.40	68.68	53.88
Wan2.1-T2V (Wan et al., 2025)	83.69	85.59	76.11	66.07	86.28	69.58	95.40	75.39	45.75
Autoregressive (Diffusion) models									
CogVideo (Hong et al., 2022)	67.01	72.06	46.83	38.18	73.40	18.11	78.20	18.24	28.24
Emu3 (Wang et al., 2024b)	80.96	84.09	68.43	59.64	86.17	44.64	77.71	68.73	37.11
NOVA (Deng et al., 2024)	80.12	80.39	79.05	59.42	92.00	77.52	95.20	77.52	54.06
SkyReel-v2 (Chen et al., 2025)	83.90	84.70	80.80	-	-	-	-	-	-
MAGI-1-distill (Teng et al., 2025)	77.92	80.98	65.68	62.43	82.37	35.08	84.20	57.75	26.28
Normalizing Flows									
STARFlow-V (Ours)	78.67	80.24	72.37	54.48	86.65	53.48	94.00	49.84	47.08
$STARFlow-V^{\dagger}$ (Ours)	79.70	80.76	75.43	59.73	80.61	56.04	98.13	76.08	48.21
STARFlow-V † (Ours, non-Causal)	79.22	80.34	74.71	58.70	81.08	54.60	98.40	73.15	49.61

Table 1 Text-to-video evaluation on VBench (Huang et al., 2024). The baseline data is from the leaderboard. Following Yang et al. (2025), we also evaluate with the official GPT-augmented prompts (Rewriter), with longer and more descriptive text inputs. † denotes results using Rewriter prompts.

Finally, we train s_{ϕ} jointly with f_{θ} at **minimal overhead**: since f_{θ} is trained by maximizing $\log p_{\theta}$, we cache the input gradients from the backward pass and reuse it as the target for s_{ϕ} .

3.3 Fast Inference

While STARFlow-V leverages parallel computation during training via causal masking, generation at inference time is carried out sequentially (one token at a time) through multiple AF blocks, which can be *extremely* computationally demanding for long video sequences. For instance, generating a 5s 480p video under 16 fps using a pre-trained 3B parameter model requires over 30 minutes, which is far from real-time performance. To enable fast inference, we introduce two strategies:

Block-wise Jacobi Iteration Rather than sampling continuous tokens strictly autoregressively, we accelerate inference by recasting inversion as solving a nonlinear fixed-point system with parallel solvers such as Jacobi iteration (Porsching, 1969; Kelley, 1995), a strategy recently used to speed up autoregressive models (Song et al., 2021; Teng et al., 2024; Liu and Qin, 2025; Zhang et al., 2025). Specifically, the inverse of Equation (2.2) can be written as the fixed-point equation

$$\boldsymbol{x} = \mu_{\theta}(\boldsymbol{x} \odot \boldsymbol{m}) + \sigma_{\theta}(\boldsymbol{x} \odot \boldsymbol{m}) \cdot \boldsymbol{z}, \tag{3.5}$$

where \boldsymbol{m} is a (self-exclusive) causal mask. This induces a triangular system that admits convergence under nonlinear Jacobi iteration (Saad, 2003): starting from an initial sequence estimate $\boldsymbol{x}^{(0)}$, iterate $\boldsymbol{x}^{(k+1)} = \mu_{\theta}(\boldsymbol{x}^{(k)} \odot \boldsymbol{m}) + \sigma_{\theta}(\boldsymbol{x}^{(k)} \odot \boldsymbol{m}) \cdot \boldsymbol{z}$ until a converge criterion is satisfied. We monitor a scale-normalized residual, $\|\boldsymbol{x}^{(k+1)} - \boldsymbol{x}^{(k)}\|_2^2 / \|\boldsymbol{x}^{(k+1)}\|_2^2 < \tau$ with $\tau = 0.001$ by default. Although the worst-case iteration count scales with sequence length (e.g., near-Markovian process), video generation exhibits strong global structure, substantially accelerating convergence in practice. The procedure is also guidance-compatible, as proposed in (Gu et al., 2025), which involves computing the guided parameters $\hat{\mu}$ and $\hat{\sigma}$ and then substituting them.

To further accelerate sampling, we adopt a block-wise Jacobi scheme in the spirit of Song et al. (2021); Liu and Qin (2025). The token sequence is partitioned into contiguous blocks of size B, which are processed sequentially across blocks but in parallel within each block. Within each block we run the Jacobi updates, while states from completed blocks are cached as context (e.g., KV cache) for subsequent blocks—analogous to standard AR inference. We also apply a video-aware initialization: for a new frame, the initial estimate $x_{n+1}^{(0)}$ is initialized from the previously converged frame $x_n^{(k)}$. Overall, we adopt block-based iteration within each AF block, yielding $\approx 15 \times$ lower inference latency relative to standard autoregressive decoding, while preserving visual fidelity.

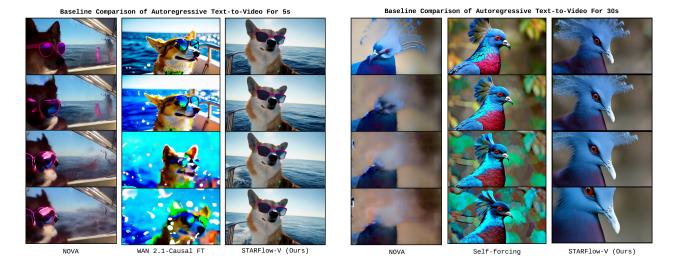


Figure 3 STARFlow-V comparison against baselines on autoregressive generation for both trained length (5s) and long-horizon generation (30s). Please refer to more video comparison in the project page.

Pipelined Decoding As described in Section 3.1, the global–local design applies standard global left-to-right autoregression in the deep block f_D , while the shallow blocks f_S traverse each frame independently. This enables a pipelined schedule (analogous to pipeline parallelism (Huang et al., 2019)): f_D runs continuously without waiting on f_S , and, in parallel, f_S threads consume f_D 's outputs, immediately refine them, and then denoise. Because f_D is typically the slowest stage, end-to-end latency is dominated by the deep block.

3.4 Versatility Across Tasks

STARFlow-V can be trained for different video generation tasks. By default, STARFlow-V is trained for text-to-video generation on large-scale text-video pairs. Without modifying the backbone, we support the following settings:

- (a) **Image-to-Video Generation.** We directly treat the first frame as observed conditioning. Owing to the invertibility, no separate encoder is required: we encode the observed frame via the flow forward to initialize the KV cache; subsequent frames are then generated.
- (b) **Video-to-Video Generation.** Given a source clip $x_{0:T}$, we treat all frames as observed conditioning and—thanks to invertibility—use the same backbone to flow-encode them and populate the KV cache. The model then autoregressively rolls out the target clip $\hat{x}_{0:T}$ under optional task cues (e.g., in/outpainting masks, edit text, camera/pose), copying through unedited regions while synthesizing edits. This mirrors our image-to-video path but operates framewise over the whole clip without a separate encoder.
- (c) **Longer Generation.** Our model generates videos far longer than those seen during training via a sliding-window (chunk-to-chunk) schedule in the deep block. After producing a latent chunk u, we warm-start the next step by rebuilding the KV cache: we re-run f_D on the last Δ latents (the overlap) and then continue autoregression to synthesize the next $N-\Delta$ latents. f_S then process the latents per frame, enabling streaming output. To mitigate boundary mismatch, we randomly drop the first frame during training to simulate restart.

4 Experiments

4.1 Experimental Setup

Datasets. We construct a diverse and high-quality collection of video datasets to train STARFlow-V. Specifically, we leverage the high-quality subset of Panda (Chen et al., 2024b) mixed with an in-house stock video dataset, with a total number of 70M text-video pairs. For all videos, we keep their raw captions, and apply a video captioner (Wang et al., 2024a) to generate a longer description to cover the details. The ratio of training

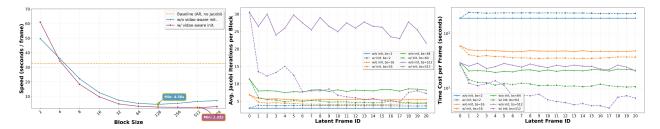


Figure 4 Comparison between speed and block size in block-wise Jacobi iteration.

using raw and synthetic captions during training is 1:9. Besides, following previous works (Lin et al., 2024), we additionally include 400M text-image pairs for joint training. To support video-to-video generation and editing, we additionally finetune the pretrained STARFlow-V on the Señorita (Zi et al., 2025), a large-scale and high-quality instruction-based video editing dataset spanning 18 well-defined editing subcategories.

Evaluation. We perform both quantitative and qualitative evaluations on STARFlow-V, and compare against baselines using VBench (Huang et al., 2024), which benchmarks text-to-video generation across 16 dimensions, including quality, semantics, temporal consistency, and spatial reasoning.

Model and Training Details. We adopt the 3D Causal VAE from WAN2.2 (Wan et al., 2025), which compresses spatial dimensions by $\times 16$ and the temporal dimension by $\times 4$ into a 48-channel latent space. We train progressively: we initialize from an image (single-frame) model, then scale to a 7B-parameter video model by increasing the deep-block capacity. For resolution, we use a curriculum from 384p to 480p while keeping the sequence length fixed at 81 frames. For the learnable denoiser, we used a 8-layer Transformer with the same channel dimension as shallow block. We include more implementation details in Appendix.

Baselines. We compare with three baselines: (i) WAN-2.1 Causal, the autoregressive variant of WAN (Wan et al., 2025) finetuned with the CausVid strategy (Yin et al., 2025); (ii) Self-Forcing (Huang et al., 2025), finetuned from WAN-2.1 Causal-FT to mitigate train—test mismatch; and (iii) NOVA(Deng et al., 2024), a native autoregressive diffusion model that does not rely on vector quantization. The original model predicts in a chunk-based fashion. For fair comparisons, we also execute results in the pure AR settings. Besides, we also report quantitative results on VBench with official scores.

4.2 Quantitative Results

Table 1 reports T2V results on VBench (Huang et al., 2024). While STARFlow-V does not yet match the strongest diffusion-based video generators, it attains performance in the same range as recent causal diffusion baselines, substantially narrowing the historical gap between NFs and diffusion models for video. To the best of our knowledge, STARFlow-V is the **first NF-based text-to-video model** to reach this level of quality, indicating that NFs can be a viable alternative when invertibility and exact likelihood (as shown in (Zhai et al., 2024)) are desired. We also include a variant trained without local constraints; its VBench scores remain very close to the causal version, indicating that enforcing causal structure does not incur a noticeable loss in perceptual quality.

4.3 Qualitative Results

T2V & I2V Tasks As illustrated in Figure 1, STARFlow-V naturally supports both T2V and I2V generation. The examples show that STARFlow-V produces temporally smooth and visually faithful sequences in both settings. Importantly, both T2V and I2V results are obtained from the *same* model without additional tuning: thanks to invertibility and causal modeling, the decoder can be reused as an encoder when a conditioning image is provided.

V2V Tasks As shown in Figure 1, STARFlow-V handles diverse V2V tasks from object-level to dense prediction within a single framework simply by changing the instruction. These results illustrate the potential of using our NF-based model for general video editing and reasoning.

Against Autoregressive Diffusion Models In Figure 3, we compare STARFlow-V with two representative autoregressive diffusion models. For the dog-with-sunglasses example, NOVA (Deng et al., 2024) exhibits gradual blurring and loss of identity, while WAN 2.1-Causal FT shows strong artifacts and color distortions. In contrast, STARFlow-V maintains clean, sharp, and temporally consistent frames, indicating stronger robustness to exposure bias. The right block of Figure 3 further shows that STARFlow-V sustains stable, coherent generations when extended to 30 seconds—well beyond its 5-second training horizon—where NOVA (Deng et al., 2024) and Self-Forcing (Huang et al., 2025) suffer from blurring, color drift, and structural deformation. We further report quantitative metrics for evaluating drifting effects across baselines and our model in the Appendix.

4.4 Ablation Study

Choice of Denoiser Figure 5 provides an ablation on the denoiser design. As shown in the top row, Decoder-finetuning (Gu et al., 2025) tends to lose temporal consistency with noticeable frame-to-frame jitter, while score-based denoising (Zhai et al., 2024) introduces bright speckle artifacts, especially in regions of large motion. The quantitative comparison (bottom) further shows that our proposed flow-score matching achieves substantially better video reconstruction under latent-space noise injection, outperforming both alternatives by a clear margin.

Hyper-parameters of Block-wise Jacobi Iteration We analyze how the block size used in the block-wise Jacobi Iteration influences the runtime of the deep block. As shown in Figure 4 (left), the runtime initially decreases as the block size increases, reflecting better utilization of intra-block parallelism, but then rises slightly again when the block size becomes too large. This trend suggests a trade-off: while larger block sizes increase parallelism, excessively large blocks requires more iterations within each block to achieve convergence.

We also examine the impact of video-aware initialization on runtime. As illustrated in Figure 4 (left), initializing the first Jacobi iteration of each frame using the converged state from the previous frame







(a) Decoder-finetuning

(b) Score-based Denoising

Method	PSNR↑	SSIM↑
No noise	32.22	0.8907
Decoder fine-tuning (Gu et al., 2025) Score-based denoising (Zhai et al., 2024) Flow-score matching (ours)	23.95 22.05 26.69	0.6403 0.6490 0.7601

Figure 5 Ablation study for the choice of denoiser. We compare video VAE reconstruction quality across denoising approaches over 1,000 random videos with large motions.

substantially reduces runtime across almost all block sizes except for small block sizes. This improvement likely stems from the strong temporal coherence present in natural videos, where neighboring frames provide effective warm starts that appear to facilitate faster iterative updates. Overall, video-aware initialization leads to observed improvements across block sizes.

We further analyze the runtime breakdown across latent frames in Figure 4 (right). Video-aware initialization yields the largest gains for large block sizes after the first frame, where convergence would otherwise require many more inner steps. Based on this observation, we adopt an asymmetric default strategy: use a medium block size (e.g., 64) for the first frame, and a larger block size (e.g., 512) for subsequent frames with videoaware initialization.

5 **Conclusion and Limitations**

We presented STARFlow-V, an end-to-end video generative model based on autoregressive normalizing flows. As shown experimentally, STARFlow-V delivers strong long-horizon coherence and fine-grained controllability across text-to-video, image-to-video and video-to-video tasks, and shows consistent gains over autoregressive diffusion baselines at 480p/81f. As a bonus, STARFlow-V can be used natively for likelihood estimation.

While the results are encouraging, there are still limitations to overcome. (1) Latency. Despite the proposed accelerated sampling, inference remains far from real time on commodity GPUs. (2) Data quality and scaling. Progress is bounded by dataset noise and bias; we do not observe a clean scaling law under current curation. (3) Non-physical generation. Due to the current model scale and available data, we still observe many unrealistic, non-physical generations (see Figure 6), such as an octopus passing through the wall of a jar and a rock spontaneously appearing beneath a goat just as it lands.

Looking forward, we see several promising directions. First, we aim to reduce generation latency, for example through more efficient sampling schedules and architectural optimizations. Second, we plan to study distillation and pruning to obtain



Figure 6 Failure cases of generation from STARFlow-V.

compact student models that retain most of the performance of the full system. Third, we will revisit dataset curation and active data selection, with a particular focus on challenging, large-motion sequences and physically grounded scenarios; this is crucial for improving physical plausibility, reducing non-physical failure cases, and enabling clearer scaling behavior at higher fidelity.

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A Derivations and Algorithms

A.1 Derivation of STARFlow-V.

(1) Why an autoregressive Gaussian model in u is a normalizing flow. Let $T_{\theta}: u \mapsto z$ be the triangular autoregressive map applied by the deep block f_D (within a frame and across frames in the global order). For token index i in that order,

$$z_i = \frac{u_i - \mu_{\theta}(u_{< i})}{\sigma_{\theta}(u_{< i})}, \quad \sigma_{\theta}(\cdot) > 0,$$
 (A.1)

with inverse

$$\mathbf{u}_i = \sigma_{\theta}(\mathbf{u}_{< i}) \mathbf{z}_i + \mu_{\theta}(\mathbf{u}_{< i}). \tag{A.2}$$

Because each z_i depends only on (u_1, \ldots, u_i) and $\sigma_{\theta} > 0$, T_{θ} is bijective and continuously differentiable. The Jacobian is lower triangular with diagonal entries $\partial z_i/\partial u_i = 1/\sigma_{\theta}(u_{< i})$, thus

$$\log \left| \det J_{T_{\theta}}(\boldsymbol{u}) \right| = -\sum_{i} \log \sigma_{\theta}(\boldsymbol{u}_{< i}). \tag{A.3}$$

With a standard normal prior $p_0(z) = \prod_i \mathcal{N}(z_i; 0, I)$,

$$\log p_D(\boldsymbol{u}) = \log p_0(T_{\theta}(\boldsymbol{u})) + \log \left| \det J_{T_{\theta}}(\boldsymbol{u}) \right| = -\frac{1}{2} \sum_{i} \boldsymbol{z}_i^2 - \sum_{i} \log \sigma_{\theta}(\boldsymbol{u}_{< i}) + \text{const}, \tag{A.4}$$

which is essentially the regression objective through maximum likelihood estimation over u. Therefore, the deep block realizes a valid normalizing flow. Composing with the shallow block gives $f_{\theta} = f_D \circ f_S$ and yields the data density in Equation (3.1).

(2) How we get the autoregressive distribution. From the global-local factorization (Equation (3.2)),

$$p_{\theta}(\boldsymbol{x}) = \prod_{n=1}^{N} p_{D}(\boldsymbol{u}_{n} \mid \boldsymbol{u}_{< n}) \left| \det J_{f_{S}}(\boldsymbol{x}_{n}) \right|, \qquad \boldsymbol{u}_{n} = f_{S}(\boldsymbol{x}_{n}).$$
(A.5)

Within a frame n, index tokens $k = 1, ..., HW \cdot D$ in raster (or block) order and we have Equation (A.4) which models p_D as Gaussian. The shallow-block contributes the additional log-det $\sum_n \log |\det J_{f_S}(\boldsymbol{x}_n)|$, forming an expressive distribution.

(3) Noise & denoising: what the model looks like. Following the noise-augmented training (§3.2), let $\tilde{x} = x + \sigma \epsilon$, $\epsilon \sim \mathcal{N}(0, I)$. The Tweedie single-step denoiser in the flow setting (Equation (3.3)) suggests the update $x \approx \tilde{x} + \sigma^2 \nabla_{\tilde{x}} \log p_{\theta}(\tilde{x})$. To avoid high-frequency artifacts and to preserve streamability, we fit a causal denoiser s_{ϕ} via flow-score matching (Equation (3.4)) and then use

$$\hat{\mathbf{x}} = \tilde{\mathbf{x}} + \sigma s_{\phi}(\tilde{\mathbf{x}}) \approx \tilde{\mathbf{x}} + \sigma^2 \nabla_{\tilde{\mathbf{x}}} \log p_{\theta}(\tilde{\mathbf{x}}), \tag{A.6}$$

where s_{ϕ} uses a block-causal mask with at most one-frame look-ahead to retain strict streamability.

```
Algorithm 1 Training STARFlow-V with noise augmentation and flow-score matching
```

```
Require: video dataset \mathcal{D}; noise level \sigma; FSM weight \lambda_{\text{den}}
            Sample mini-batch \boldsymbol{x} \sim \mathcal{D} and noise \boldsymbol{\epsilon} \sim \mathcal{N}(0, I)
 2:
            Noise-augment: \tilde{x} \leftarrow x + \sigma \epsilon
                                                                                                                                                                     ⊳ as in §3.2
 3:
                                                                                                         ▷ alternating masked AF blocks, within-frame
            Shallow forward: u \leftarrow f_S(\tilde{x})
 4:
            Deep forward: z \leftarrow f_D(u)
                                                                                                               ▷ causal Transformer AF over global order
 5:
            Standard NF NLL: \mathcal{L}_{\text{NLL}}(\theta) \leftarrow -\left[\log p_0(z) + \log |\det J_{f_D}(u)| |\det J_{f_S}(\tilde{x})|\right]
 6:
            Score target (stop-grad): \mathbf{t} \leftarrow \sigma \nabla_{\tilde{\boldsymbol{x}}} \log p_{\theta}(\tilde{\boldsymbol{x}})
                                                                                                                    \triangleright reuse backward pass of \mathcal{L}_{NLL}; detach
 7:
            Flow-score Matching: \mathcal{L}_{\text{FSM}}(\phi) \leftarrow ||s_{\phi}(\tilde{x}) - \mathbf{t}||_2^2
 8:
            Total loss: \mathcal{L} \leftarrow \mathcal{L}_{NLL}(\theta) + \lambda_{den} \mathcal{L}_{FSM}(\phi)
 9:
            Update: (\theta, \phi) \leftarrow (\theta, \phi) - \eta \nabla \mathcal{L}
10:
11: until convergence
```

Algorithm 2 Autoregressive sampling $(z \rightarrow u \rightarrow x)$

```
Require: length N (frames or tokens), base prior p_0(z) = \mathcal{N}(0, I), shallow inverse f_S^{-1}, deep inverse f_D^{-1}, token order \prec
```

- 1: Sample $z \sim \mathcal{N}(0, I)$ with the target shape
- 2: Initialize an empty latent sequence \boldsymbol{u}
- 3: for each element i in global order \prec do \triangleright causal AR over frames and within-frame tokens
- 4: Compute (μ_i, σ_i) : $(\mu_i, \sigma_i) \leftarrow f_D(\mathbf{u}_{< i})$
- 5: Invert deep at position $i: \mathbf{u}_i \leftarrow \sigma_i \mathbf{z}_i + \mu_i$ $\triangleright f_D^{-1}$, triangular
- 6: end for
- 7: Invert shallow block: $\boldsymbol{x} \leftarrow f_S^{-1}(\boldsymbol{u})$
- 8: (One-step corrector) $x \leftarrow x + \sigma_{\text{test}} s_{\phi}(x)$
- 9: $\mathbf{return} \ x$

A.2 Training

Algorithm 1 shows the training algorithm of STARFlow-V for both the flow and the learnable denoiser.

A.3 Inference

Remarks. (i) When the deep map is sufficiently contractive in u (e.g., via scale clamping), the Jacobi iteration converges rapidly and enables wide parallelism within each block B. (ii) A common choice for \mathcal{B} is to use spatial tiles per frame (no intra-tile dependencies) or even/odd raster groups, preserving the block-causal mask used in training.

B Implementation Details

B.1 Architecture Design

	3B	7B
Params	$\sim 3B$	~7B
f_D width	3072	4096
f_S		masked AF; width d_S , depth L_S)
Denoiser s_{ϕ}	8-layer Tr	cansformer, block-causal mask
Init	from scratch	finetune from 3B

Table 2 Minimal comparison. Only f_D width differs; f_S and s_ϕ are unchanged.

3B. Same size as STARFlow but for video. The deep block f_D uses width 3072 (depth L_D , heads H_D). The

Algorithm 3 Jacobi-style parallel inversion of the deep autoregressive block

```
Require: base latent z; initial guess u^{(0)} (e.g., zeros); block partition \mathcal{B} = \{B_1, \dots, B_J\} (non-overlapping,
       block-causal, |B_i| = 4|B_1| for all block j > 1); max iters T; Frame size F; tolerance \tau
  1: for j = 1, 2, ..., J do
             [a,b] \leftarrow B_j
                                                                                                                                                      \triangleright indices of the j-th block
  2:
             if j = 1 and a > F then
  3:
                   Initialize \boldsymbol{u}_{a:b} \leftarrow \boldsymbol{u}_{a:b}^{(0)}
                                                                                                                                                           ▷ random initialization
  4:
  5:
                   Initialize u_{a:b} \leftarrow u_{a-F:b-F}
                                                                                                                                            ▷ initialization from past frame
  6:
             end if
  7:
             repeat
  8:
                   t \leftarrow t + 1
  9:
                   for all i \in B_i in parallel do
10:
                         (\mu_i^{(t)}, \sigma_i^{(t)}) \leftarrow f_D(\boldsymbol{u}_{< i}^{(t)}) \ \boldsymbol{u}_i^{(t+1)} \leftarrow \sigma_i^{(t)} \, \boldsymbol{z}_i + \mu_i^{(t)}
11:
12:
             \begin{array}{l} \mathbf{end} \ \mathbf{for} \\ \mathbf{until} \ \frac{\| \boldsymbol{u}^{(t+1)} - \boldsymbol{u}^{(t)} \|_2}{\| \boldsymbol{u}^{(t)} \|_2 + \varepsilon} \leq \tau \ \mathbf{or} \ t = T \end{array}
13:
14:
             oldsymbol{u}_{a:b} \leftarrow oldsymbol{u}_{a:b}^{(t)}
15:
16: end for
17: Shallow inverse: x \leftarrow f_S^{-1}(u^{(t+1)})
18: (One-step corrector) x \leftarrow x + \sigma_{\text{test}} s_{\phi}(x)
19: return x
```

shallow stack f_S (alternating masked affine flows) and the denoiser s_{ϕ} (8-layer Transformer with block-causal mask) follow the standard design.

7B. Initialized from the 3B checkpoint and *only* widens the deep block f_D channels from 3072 to 4096. The shallow stack f_S and denoiser s_{ϕ} remain identical (same depths, heads, and widths).

B.2 Training Details

STARFlow-V is trained on 96 H100 GPUs using approximately 20 million videos. In all the experiments, we share the following training configuration for our proposed STARFlow-V.

```
training config:
   batch_size=96
   optimizer='AdamW'
   adam_beta1=0.9
   adam_beta2=0.95
   adam_eps=1e-8
   learning_rate=5e-5
   min_learning_rate=1e-6
   learning_rate_schedule=cosine
   weight_decay=1e-4
   mixed_precision_training=bf16
```

Progressive Video Training We adopt a progressive multi-stage training paradigm that gradually increases model size, resolution, and temporal horizon for stable and effective optimization.

- 3B Text-to-Image Training: We initialize a 3B text-to-image model from the pretrained StarFlow (Gu et al., 2025), establishing a strong visual-textual backbone before introducing temporal modeling.
- 3B Image-Video Joint Training (384P, 45 frames): The 3B model is then jointly trained on low-resolution images and videos at 384P. Each training clip contains 45 frames sampled at 16 fps, enabling the model to acquire short-term temporal dynamics.

Algorithm 4 Streaming long-sequence generation via re-encode with forward

```
Require: target length T (frames), window size W (W \ll T); deep inverse f_D^{-1}; shallow inverse f_S^{-1}; shallow
     forward f_S; deep forward f_D; prior p_0(z)
 1: Initialize caches KV \leftarrow \emptyset, latent buffer U \leftarrow \emptyset
     for t = 1 to T do
          Sample base: z_t \sim \mathcal{N}(0, I) for the next frame (or token block)
 3:
          Deep inverse: using cached state, compute u_t \leftarrow f_D^{-1}(z_t; \mathsf{KV}) and update the \mathsf{KV} cache.
 4:
          Shallow inverse: \boldsymbol{x}_t \leftarrow f_S^{-1}(\boldsymbol{u}_t)
 5:
 6:
          Re-encode (forward): \hat{\boldsymbol{u}}_t \leftarrow f_S(\boldsymbol{x}_t)
 7:
                                                                                      \triangleright brings the produced frame back to U-space
          Update deep state: run f_D forward on \hat{\boldsymbol{u}}_t to refresh KV (no sampling): \underline{\phantom{}} \leftarrow f_D(\hat{\boldsymbol{u}}_t; \mathsf{KV})
 8:
          Maintain sliding window: push \hat{u}_t into buffer U; if |U| > W pop the oldest
10: end for
11: return \{\boldsymbol{x}_t\}_{t=1}^T
```

- 7B Image-Video Joint Training (384P, 81 frames): We expand the model to 7B parameters and continue joint training at 384P, doubling the temporal horizon from 45 to 81 frames to strengthen long-range temporal reasoning.
- 7B Image-Video Joint Training (480P, 81 frames): Finally, we train the 7B model on higher-resolution 480P images and videos while maintaining the 81-frame temporal window.

Mixed-Resolution Training STARFlow-V is designed to support *mixed-resolution* inputs, allowing each frame to retain its native aspect ratio and spatial resolution. Similar to Gu et al. (2025), we assign each video sequence to one of nine predefined aspect-ratio bins, since all frames within a video share the same ratio. The pre-defined bins are 21:9, 16:9, 3:2, 5:4, 1:1, 4:5, 2:3, 9:16, and 9:21. To make the model explicitly aware of these visual formats, we incorporate both the fps and aspect-ratio tag into the text caption:

```
A video with {fps} fps:
{original_caption}
in a {aspect_ratio} aspect ratio.
```

Gradient Control We monitor the gradient norm throughout training to ensure stability. Specifically, to prevent gradient explosion, we enable gradient skipping after the first 100 steps: if the gradient norm exceeds a threshold of 1, the update for that step is skipped. This adaptive strategy stabilizes early training while maintaining convergence efficiency later on.

B.3 Baseline Details

WAN-2.1 Causal-FT is the autoregressive variant of WAN (Wan et al., 2025). Specifically, we adopt Wan2.1-T2V-1.3B, a Flow Matching-based model, as the base model. Following the CausVid initialization strategy (Yin et al., 2025), the base model is fine-tuned with causal attention masking on 16k ODE solution pairs generated from the model itself. In practice, we leverage the ODE initialization checkpoint released with the official Self-Forcing (Huang et al., 2025) repository, which corresponds exactly to the configuration of our WAN-2.1 Causal-FT setup.

NOVA AR (Deng et al., 2024) is an autoregressive video generator that does not rely on vector quantization. It reformulates video generation as non-quantized autoregressive modeling that performs temporal frame-by-frame prediction while generating spatial token sets within each frame in a flexible, set-by-set manner. To support autoregressive modeling with continuous tokens, NOVA leverages a lightweight diffusion head that models the distribution of each continuous token (Li et al., 2024). In this work, we directly compare the pure AR version of NOVA, where the model predicts each latent frame with diffusion for a fair comparison.

Model	Total	Quality	Semantic	Aesthetic	Object	Human	Spatial	Scene
Autoregressive (Diffusion) models								
NOVA AR† (Deng et al., 2024)	75.31	77.46	66.70	56.04	79.68	94.20	66.07	47.83
WAN 2.1-Causal FT†	74.96	77.41	65.15	56.04	76.51	94.20	53.25	47.83
Normalizing Flows								
STARFlow-V† (Ours)	79.70	80.76	75.43	59.73	80.61	98.13	76.08	48.21

Table 3 Performance comparison of autoregressive video generation models on VBench (Huang et al., 2024). Following Yang et al. (2025), we evaluate with the official GPT-augmented prompts (noted as †)

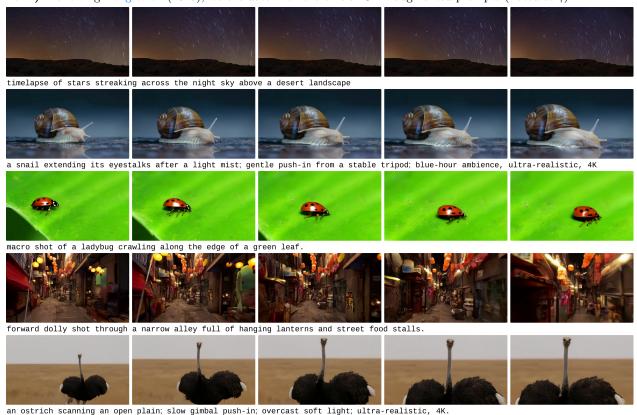


Figure 7 Generated samples from STARFlow-V given text prompts. All videos are at 480p 16fps and 5s.

C Additional Experimental Details and Results

C.1 Quantitative Comparison with Autoregressive Diffusion baselines

To evaluate the robustness of video generation under autoregressive generation, we compare STARFlow-V with autoregressive diffusion models, including NOVA AR (Deng et al., 2024) and WAN 2.1-Causal FT. Here, NOVA AR refers to the fully autoregressive video generation variant which is different from the reported in the official paper. Table 3 compares these models across a diverse set of evaluation dimensions defined in VBench (Huang et al., 2024). As shown in Table 3, STARFlow-V substantially outperforms the autoregressive diffusion baselines across all dimensions. Both NOVA AR and WAN 2.1-Causal FT exhibit clear signs of autoregressive degradation in their generated videos. Specifically, NOVA AR suffers from pronounced error accumulation, leading to increasing blur and content collapse as the video progresses. And WAN 2.1-Causal FT produces noticeable temporal inconsistency and flickering throughout the video. These failure modes are reflected in their lower scores, underscoring the difficulty of maintaining robustness in autoregressive video generation. And it further highlights the strength of our approach.

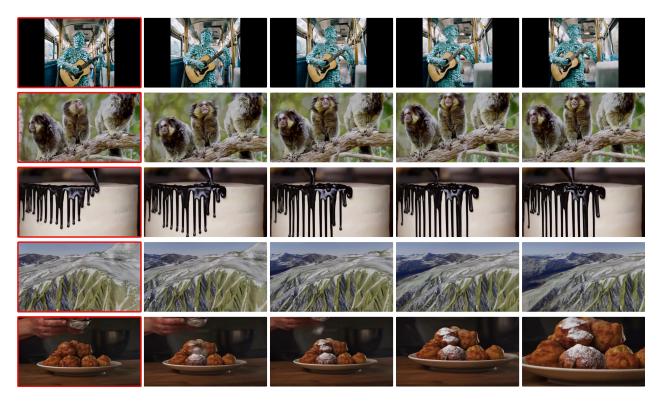


Figure 8 Generated samples from STARFlow-V given the first frame. All videos are at 480p 16fps and 5s.

C.2 Video-to-Video Generation

To support video-to-video generation and editing, we additionally finetune the pretrained STARFlow-V (7B, 384P, 81 frames) on the Señorita (Zi et al., 2025), a large-scale and high-quality instruction-based video editing dataset spanning 18 well-defined editing subcategories. Each training sample in Señorita consists of a 33-frame input video paired with a 33-frame edited target video. The model is also trained on videos with 16fps. This finetuning stage equips STARFlow-V with precise editing capabilities while preserving temporal coherence and motion consistency. During finetuning, we concatenate the input and target videos along the temporal dimension to form a single training sequence.

C.3 Additional Samples

We show additional samples at Figures 7 to 9. Besides, we provide more video generation comparison in our official codebase at https://github.com/apple/ml-starflow.

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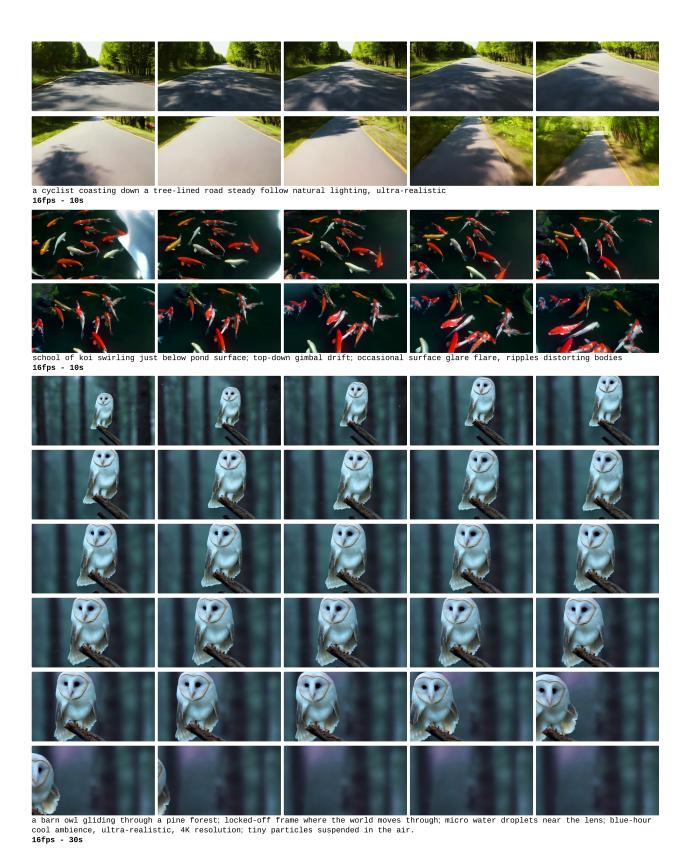


Figure 9 Generated samples from STARFlow-V given text prompts and extended with overlapping frames. For each segment, we generate 21 latent frames with 4 latent frames in overlap. Both videos are at 480p 16fps.